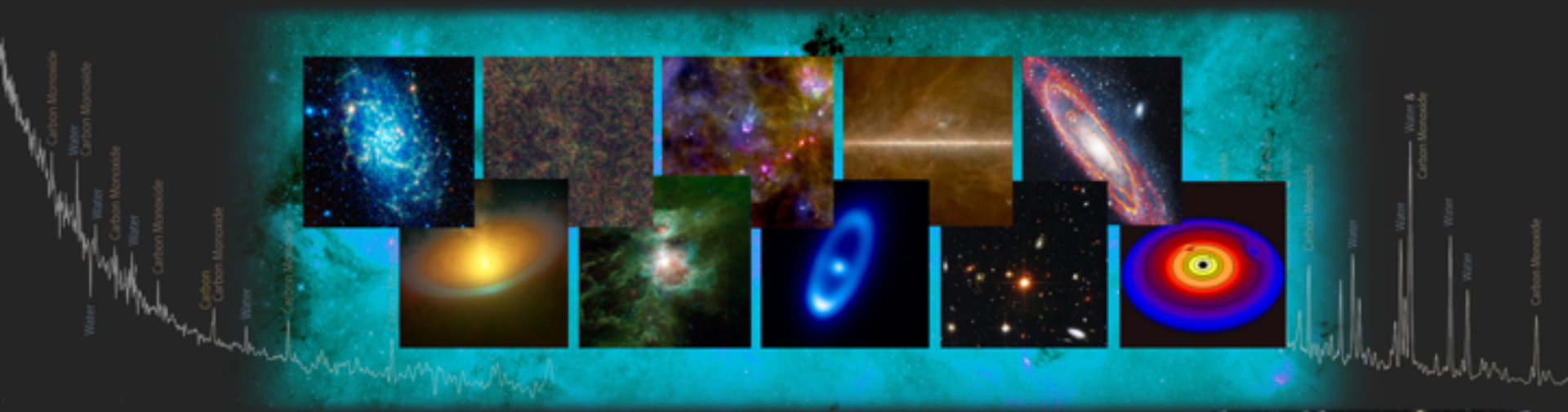
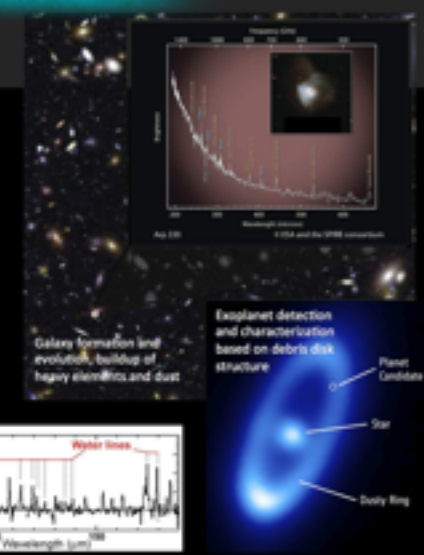


# Satellite Meeting 2: Mission Concepts and Measurement Requirements for a Future Far-Infrared Space Mission



Dave Leisawitz  
NASA/GSFC  
Pathways 2015  
17 July 2015



# Participants



- Bill Danchi, NASA Goddard Space Flight Center, US
- Steve Ertel, ESO, Chile
- Antonio Garcia Munoz, ESA/ESTEC, The Netherlands
- Jane Greaves, University of St. Andrews, UK
- Satoshi Itoh, Osaka University, Japan
- Grant Kennedy, University of Cambridge, UK
- Dave Leisawitz, NASA Goddard Space Flight Center, US
- Laura McKemmish, UCL, UK
- Mike Meyer, ETH Zurich, Switzerland
- Eugenio Schisano, IAPS/INAF, Italy
- Feng Tian, Tsinghua University, China
- John Trauger, Caltech-JPL, US
- Steve Unwin, Caltech-JPL, US
- Joachim Wiegert, Chalmers Univ. of Technology, Sweden
- Siyi Xu, ESO, Germany
- Hans Zinnecker, Universität Stuttgart and NASA-Ames, US

**16+ people**  
**10 nations**

# Purpose of the meeting



**To serve as an opportunity for the participants to ...**

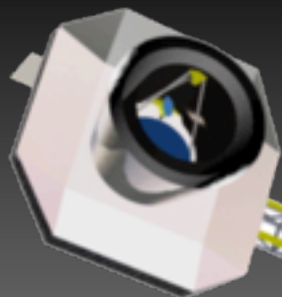
- **learn about concepts for future far-infrared space missions, and learn what is technically feasible and affordable in a mission that could begin in the 2020s; and**
- **discuss measurement requirements for the mission, including angular resolution, spectral resolution, and sensitivity.**

# Mission concepts



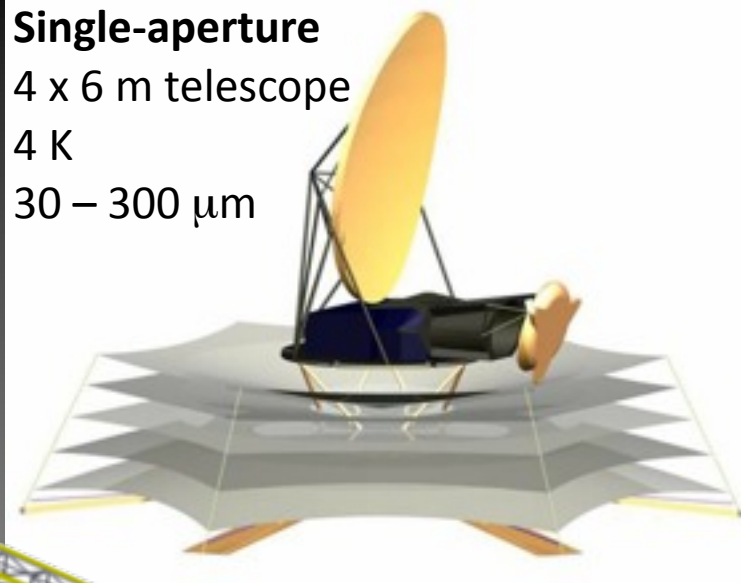
## Far-IR Interferometer

1 m class telescopes  
4 K  
25 – 400  $\mu\text{m}$



## Single-aperture

4 x 6 m telescope  
4 K  
30 – 300  $\mu\text{m}$



## SPICA

2.5 m telescope  
8 K  
17 – 230  $\mu\text{m}$

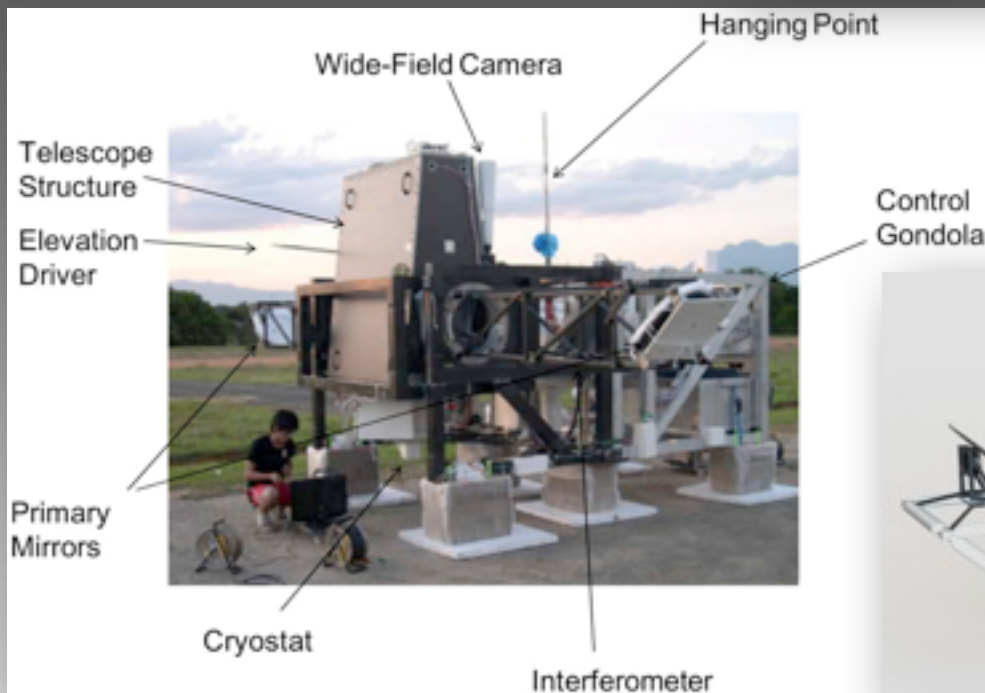
- Different measurement capabilities
- Different costs
- All have spectroscopic capability

## SOFIA



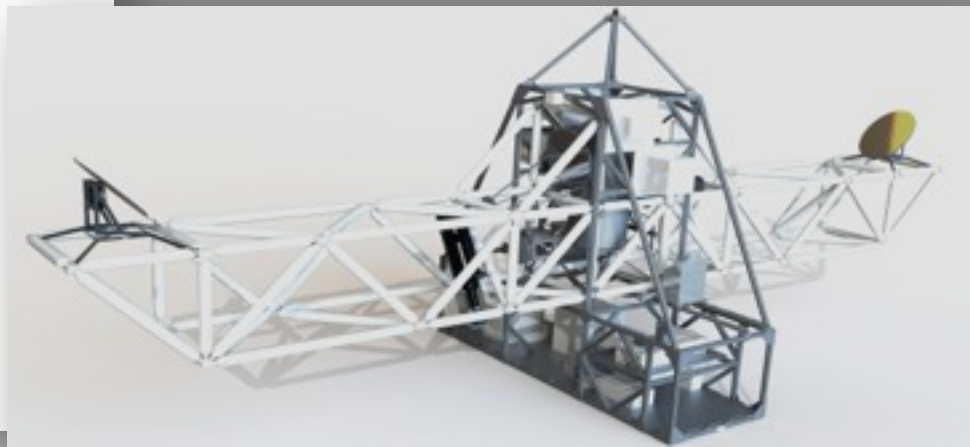
## FITE

H. Shibai, Osaka University, PI  
8 m Fizeau far-IR interferometer



## BETTII

S. Rinehart, NASA GSFC, PI  
8 m "double Fourier" interferometer for spatially resolved spectroscopy

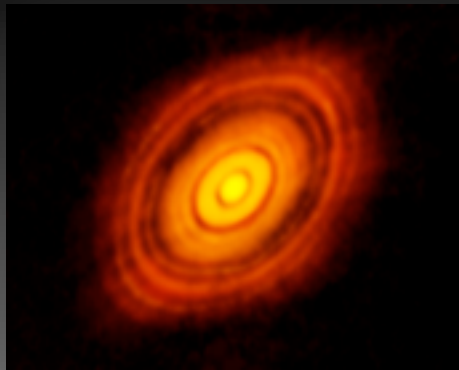


# Single-Aperture Apps



- Most disks would not be resolved, but many would be detected, and one could:
  - measure the total molecular gas ( $H_2$ , HD) and water (gas phase and ice) in disks, from protostellar to protoplanetary to transitional to debris disks.
  - put the solar system in context by measuring the debris disk luminosity function down to solar system level
  - characterize the contents of debris disks (grain size distribution and composition)

# Killer (?) Apps



ALMA, HL Tau

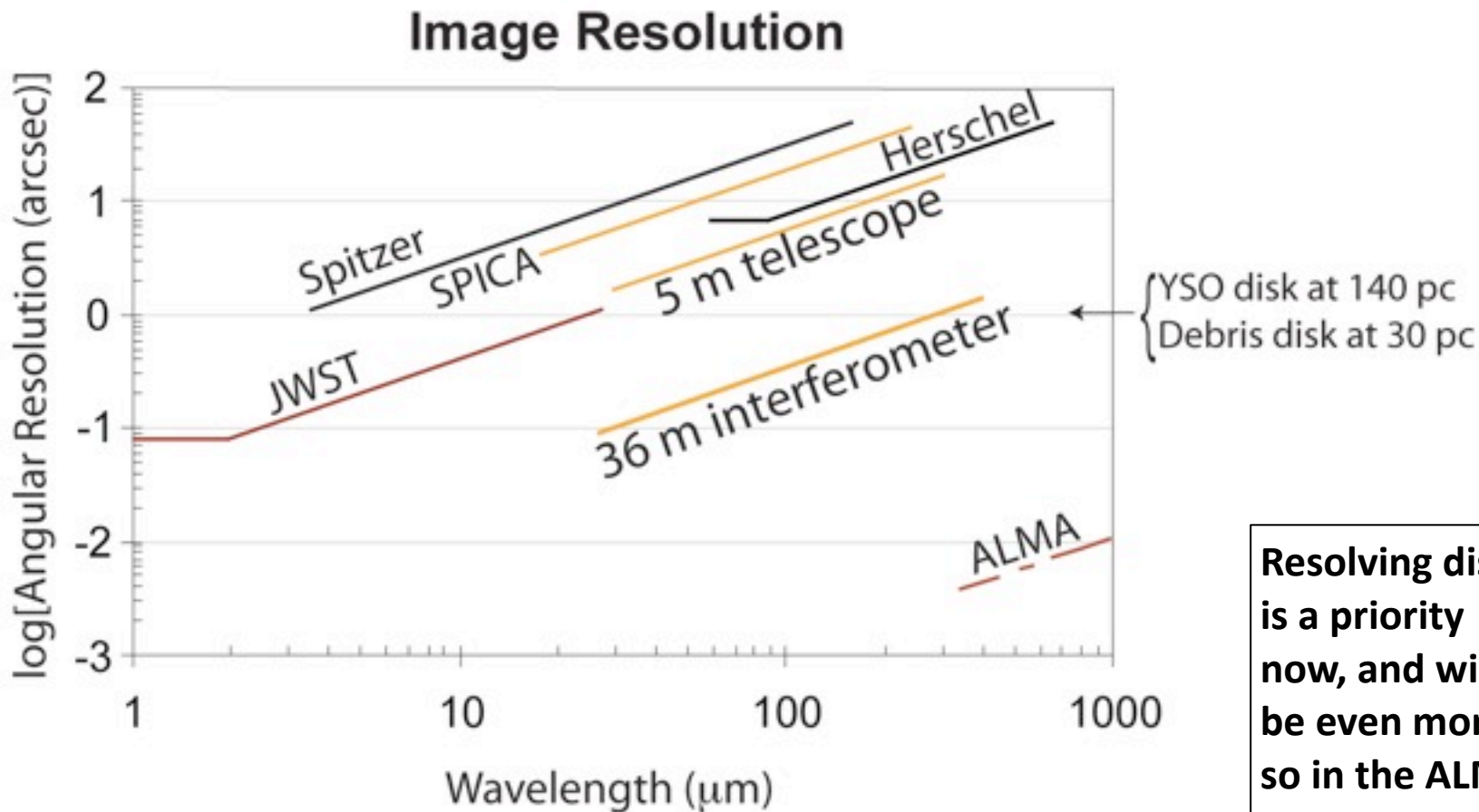
Image protoplanetary disks and measure the distributions of  $H_2$ , HD, water vapor, ice, and dust to learn how the conditions for habitability arise during the planet formation process.

Image structures in a large number of debris disks to find and characterize exoplanets through their interactions with the disks.



Herschel, Fomalhaut

# Desired measurement capabilities of the Satellite 2 meeting group tend to favor interferometry



**Resolving disks is a priority now, and will be even more so in the ALMA and JWST era.**



# Complex Landscape



- The NASA Astrophysics Roadmap calls for a “FIR Surveyor” to deliver “crucial science” in the Formative Era (2020s), where it would serve as “a logical starting point” and “a training ground” for more ambitious shorter-wavelength interferometers in the Visionary Era (2030s and beyond).
- Members of the far-IR community met last month and voted for a 4 x 6 m single-aperture telescope
  - “Galaxy evolution” (extragalactic spectroscopy) was the dominant science driver
  - Mission cost ostensibly <\$2B
  - Chief objective was to convince NASA to invest in an STDT study of the “FIR Surveyor”
- This and other STDTs will develop mission concepts that are considered most likely to succeed in the Decadal Survey.